

## Calibration of Earth observation hyperspectral instruments - some examples

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### Outline



apportionment

instrument design - calibration - L0 to L1b processing

LO performance verification and calibration

- calibration sources
- calibration parameters
- lessons learnt

#### space segment compliance with instrument Level 1b requirements at BOL and EOL Instrument design **Opto-mechanical** Thermal Electrical I/FData processing Calibration LO-L1b L0/L1bATBD optimisation considering L1b instrument performance at L1b **On-ground** accuracy budgets and programmatic aspects

## Apportionment - delicate balance







#### **Calibration sources**



#### **EXTERNAL**

Fraunhofer lines atmospheric emission lines Sun moon stars OGSE's vicarious

#### INTERNAL

. . .

WLS PRNU, radiometric stability, degradation LED dead/bad pixel, pixel gain, pixel linearity diffuser nominal/ redundant spectral line source lamp/ laser

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#### **S4/UVN** instrument





ZUVN

## **S4/UVN** calibration sources





### **Calibration parameters**



- radiometric calibration Earth spectral radiance, Sun spectral irradiance and Earth reflectance (radiance/irradiance)
- → spectral calibration
  - wavelength scale for uniform and non-uniform ground scenes
  - Instrument Spectral Response Function (ISRF)
  - optical bench temperature (gradient) dependencies
- → spectral/ spatial straylight
- → electronic and detector calibration parameters
- → geometric parameters, co-registration, Image Navigation and Registration (INR), geolocation
- → polarisation correction, verification at level 0 that instrument is
  sufficiently insensitive to incident polarisation or polarisation calibration
- → at level 0: no spectral features in instrument response that interfere with atmospheric gas absorptions

### **Radiometric calibration parameters**



- 1. absolute Earth spectral radiance
- 2. absolute Sun spectral irradiance

#### 3. Earth viewing angle dependency

North-South on detectors and scan mirror (OMI and OMPS have no scan mirror)

- 4. Sun viewing angle dependency
- 5. absolute Earth **reflectance**: Earth radiance/ solar irradiance, using dedicated sources optimised for this parameter

In orbit, relative radiometric **degradation** monitoring and quantification primarily with **Sun irradiance** measurements, but also with **WLS** and **LED**, Earth **radiance** and **moon** measurements.

## Absolute radiometric radiance calibration - on ground



Absolute radiometric radiance calibration measurements using calibrated sources (FEL lamp, integrating sphere) and **radiance angle dependency calibration** measurements under flight-representative thermal-vacuum conditions



## FEL lamp as part of OGSE





## Diffuser as part of OGSE





## Integrating sphere as part of OGSE





#### ... and do not forget Theo(dolite)!



# Predecessor atmospheric (hyperspectral) instruments



- 1. TOMS (NASA)
- 2. GOME
- 3. Sciamachy (ENVISAT)
- 4. GOME-2 (Metop)
- 5. OMI (Aura)
- 6. OMPS (NPP, NASA)
- 7. ...and many more...

#### SCIAMACHY (Envisat)



Azimuth Scanner

Design by TNO

- → channel 2/8 breadboard calibrated in 2-3 months
- → 5 tests in TV chamber (OPTEC) were necessary on the PFM lasting between 6/9 months
- → SCIAMACHY not possible to rotate in vacuum
- many new effects, as e.g. "spectral features" effect has been discovered on SCIAMACHY instrument level, resulting in late change of solar diffuser (additional diffuser)

Note: L0, Cal phase separation revealed to be crucial, because performance measurements revealed issues that had to be solved on hardware before calibration could start (e.g. channel out of focus, PMD electronic defect) and could have potentially invalidated some of the calibration measurements if discovered too late

## SCIAMACHY OGSE example





## GOME-2 (Metop)



- → redo of GOME (first instrument of its class in Europe)
- → using an EQM, calibrated in dry-run campaign
- → still issues arisen during flight model calibration leading to a re-ship of the instrument back to the calibration facility (radiometric calibration deficiencies not understood for a long time)







## OMI (AURA)

- first instrument of this kind with a 2dimensional detector
- → using an OMI BB model (BBM), resembling OMI PFM as much as possible and performing the full performance verification & calibration campaign, using all OGSE and performing full analysis of the measurements data. Duration of about 3 months (measurement time).
- → it was not possible to rotate OMI in vacuum
- → PFM on-ground campaign was about 8 months, including a separation of performance verification and calibration phase (TV).

<u>Note:</u> separation revealed to be crucial because some of the calibration measurements would have been invalidated by some of the performance results.



### Lessons learnt from instruments



#### On ground 'Test as you fly'

→ i.e. thermal vacuum environment with flight representative conditions for pressure, temperatures for detectors and 'optical bench'

#### Consider EQM prior to flight model

→ calibrated in dry-run campaign, elimination of non-conformances

#### On ground and in flight

- → several parameters can only be measured with sufficient accuracy on ground, e.g. polarisation, straylight and ISRF
- → clear split between L0 performance verification and calibration

#### Instrument built as designed?

→ L0 performance verification

#### Calibration

→ retrieve calibration key parameters for L0 to L1b processing

#### Lessons learnt



Of paramount importance is that at the start of a programme,

- scientists,
- calibration staff,
- instrument operations staff and
- system and software developers

closely work together to identify the

- measurement types,
- the frequency of scheduling,
- the required corrections steps that are applied and
- data products



Thank you for your attention!

European Space Agency